

# La deuxième loi de Kepler – avec XINT et PSTricks

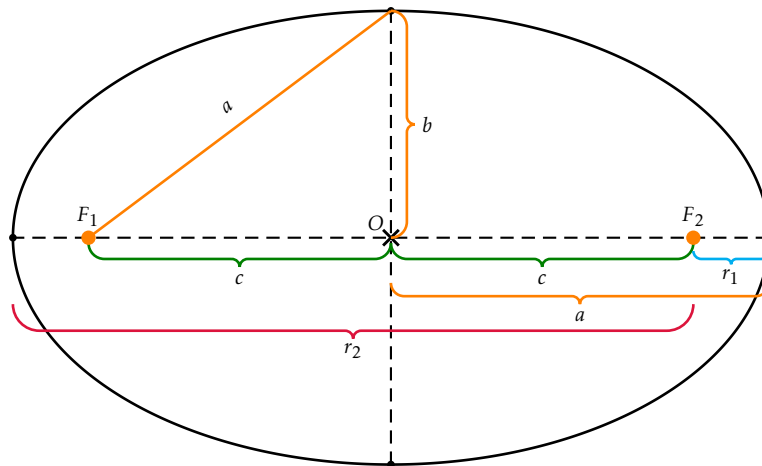
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## 1 Étude mathématique des ellipses

Dans le dessin suivant on présente la syntaxe des paramètres d'une ellipse.

- $O$  centre de l'ellipse
- $a$  demi-grand axe
- $b$  demi-petit axe
- $c$  la distance séparant le centre de l'ellipse de l'un des foyers
- $e$  l'excentricité de l'ellipse :  $e = \frac{c}{a}$

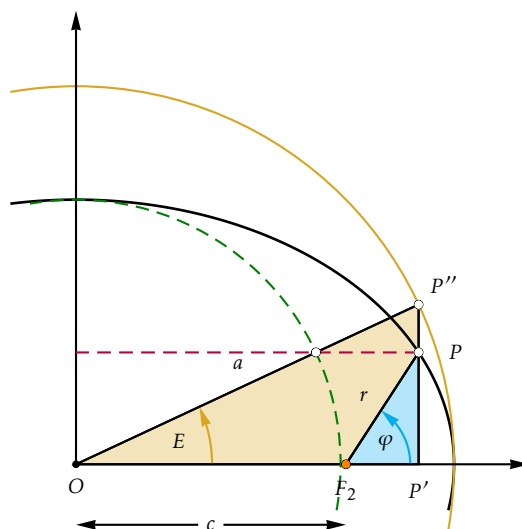


Avec le théorème de Pythagore :

$$c^2 = a^2 - b^2$$

## 2 Deuxième loi de Kepler – la théorie

La transformation de l'angle  $E$  (angle au centre de l'ellipse) avec l'angle  $\varphi$  (l'angle de sommet le foyer de l'ellipse – l'angle polaire).



Pour l'angle polaire  $\varphi$  :

$$\cos(\varphi) = \frac{a \cdot \cos(E) - c}{r} = \frac{a \cdot \cos(E) - c}{a - c \cdot \cos(E)} \quad (1)$$

$$\sin(\varphi) = \frac{b \cdot \sin(E)}{r} = \frac{\sqrt{a^2 - c^2} \cdot \sin(E)}{a - c \cdot \cos(E)} \quad (2)$$

La formule de transformation entre les angles  $E$  (origine) et  $\varphi$  (foyer  $F_2$ ) :

$$\tan\left(\frac{\varphi}{2}\right) = \sqrt{\frac{a+c}{a-c}} \cdot \tan\left(\frac{E}{2}\right)$$

En prenant la formule (2) et en dérivant par rapport à  $t$  :

$$\begin{aligned} \frac{d}{dt}(\sin(\varphi)) &= \cos(\varphi) \cdot \dot{\varphi} \\ &= \frac{b \cos(E) \cdot \dot{E}(a - c \cos(E)) - c \sin(E) \cdot \dot{E} b \sin(E)}{(a - c \cos(E))^2} \\ &= \dots \\ &= b \dot{E} \cdot \frac{a \cos(E) - c}{(a - c \cos(E))^2} \end{aligned}$$

On divise par  $\cos(\varphi)$ , alors

$$\dot{\varphi} = \frac{d\varphi}{dt} = \frac{b \dot{E}}{a - c \cos(E)} = \frac{b \dot{E}}{r} \quad (3)$$

La deuxième loi de Kepler (*Loi des aires*) dit :

*Le rayon-vecteur reliant une planète au Soleil balaie des aires égales en des temps égaux.*

L'aire doit être constante :

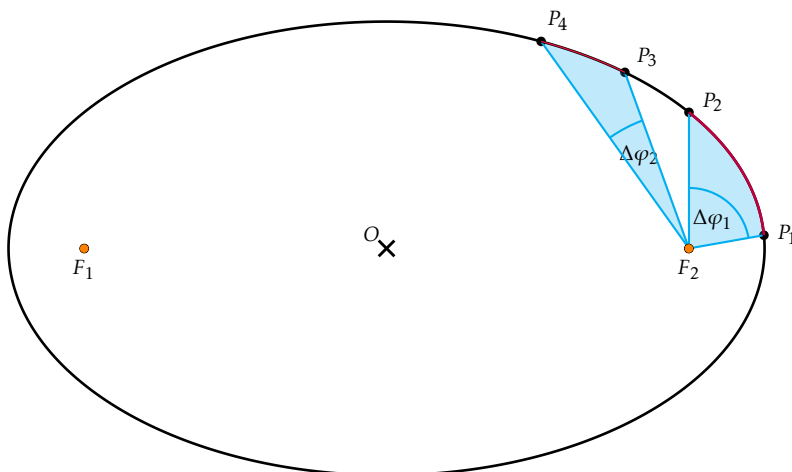
$$r^2 \cdot \dot{\varphi} = \text{cste.} = C$$

L'équation (3) résolue pour  $\dot{E}$  :

$$\dot{E} = \frac{dE}{dt} = \frac{r^2 \dot{\varphi}}{br} = \frac{C}{br} = \frac{C}{b(a - c \cos(E))}$$

Par integration

$$\begin{aligned} \int dt &= \int \frac{b}{C} (a - c \cos(E)) dE \\ t = t(E) : t &= \frac{b}{C} (aE - c \sin(E)) \end{aligned}$$



### 3 Commandes

On définit les paramètres de l'ellipse :

```
\xintdefvar Ox, Oy := 7, 2; % center of the ellipse
\xintdefvar a_e1, b_e1 := 5, 3; % demi axes
\xintdefvar phi_1, phi_2, phi_3 := 10, 90, 110; % polar angles

\xintdefvar e_e1 := sqrt(a_e1^2-b_e1^2); % distance focus center
\xintdefvar F_1x, F_1y, F_2x, F_2y := Ox-e_e1, Oy, Ox+e_e1, Oy; % coordinates of the foci
\xintdefvar eps_e1 := e_e1/a_e1; % eccentricity
```

On définit la fonction de l'ellipse :

```
\xintdeffloatfunc trans_phi_M(x) := ifsgn([Ellipse(x)][0]-Ox,
    atand([Ellipse(x)][1]-Oy)/([Ellipse(x)][0]-Ox)+180,
    0,
    atand([Ellipse(x)][1]-Oy)/([Ellipse(x)][0]-Ox)
    );
\xintdeffloatfunc trans_M_phi(x) := ifsgn([Ellipse_M(x)][0]-F_2x,
    atand([Ellipse_M(x)][1]-F_2y)/([Ellipse_M(x)][0]-F_2x)+180,
    0,
    atand([Ellipse_M(x)][1]-F_2y)/([Ellipse_M(x)][0]-F_2x)
    );
\xintdeffloatfunc phi_Pol_to_E(x) := if(x=180,
    180,
    if(x>180,
        2*atand(sqrt((a_e1-e_e1)/(a_e1+e_e1))*tand(x/2))+360,
        2*atand(sqrt((a_e1-e_e1)/(a_e1+e_e1))*tand(x/2))
    )
    );
\xintdeffloatfunc E_to_phi_Pol(x) := if(x=180,
    180,
    if(x>180,
        2*atand(sqrt((a_e1+e_e1)/(a_e1-e_e1))*tand(x/2))+360,
        2*atand(sqrt((a_e1+e_e1)/(a_e1-e_e1))*tand(x/2))
    )
    );
```

Les formules pour la transformation des angles  $E$  en  $\varphi$  et vice versa :

```
\xintdefunc trans_phi_M(x) := ifsgn([Ellipse(x)][0]-Ox,
    atand([Ellipse(x)][1]-Oy)/([Ellipse(x)][0]-Ox)+180,
    0,
    atand([Ellipse(x)][1]-Oy)/([Ellipse(x)][0]-Ox)
    );
\xintdefunc trans_M_phi(x) := ifsgn([Ellipse_M(x)][0]-F_2x,
    atand([Ellipse_M(x)][1]-F_2y)/([Ellipse_M(x)][0]-F_2x)+180,
    0,
    atand([Ellipse_M(x)][1]-F_2y)/([Ellipse_M(x)][0]-F_2x)
    );
% phi --> E
\xintdeffunc phi_Pol_to_E(x) := if(x=180,
    180,
    if(x>180,
        2*atand(sqrt((a_e1-e_e1)/(a_e1+e_e1))*tand(x/2))+360,
        2*atand(sqrt((a_e1-e_e1)/(a_e1+e_e1))*tand(x/2))
    )
    );
% E --> phi
\xintdeffunc E_to_phi_Pol(x) := if(x=180,
    180,
    if(x>180,
        2*atand(sqrt((a_e1+e_e1)/(a_e1-e_e1))*tand(x/2))+360,
        2*atand(sqrt((a_e1+e_e1)/(a_e1-e_e1))*tand(x/2))
    )
    );
```

Integration et la méthode de Newton pour calculer  $\varphi_4$  :

```
%% INTEGRATION -- ATTENTION -- calculate in RAD not DEGREES
% ---- ATTENTION --- phi_1, phi_2, phi_3 ARE NOW HARD CODED IN THE FUNCTIONS HERE ----
% NEWTON: gets root by x <-- x - delta, delta = f(x)/f'(x)
% so let's simplify f in order for computations to have been done already
% *oneDegree converts from degrees to radians
\xintdeffloatvar phiE_1, phiE_2, phiE_3 := seq(phi_Pol_to_E(x), x = phi_1, phi_2, phi_3);
\xintdeffloatfunc f(x, a, e) := a*x-e*sin(x)+e*(sind(phiE_2)+sind(phiE_3)-sind(phiE_1))-a*(phiE_2+phiE_3-phiE_1)*oneDegree;
\xintdeffloatfunc fD(x, a, e) := a-e*cos(x); % derivative of f(x)
```

```

% multiplication by oneRadian converts from Radians to Degrees
\xintdeffloatvar phiE := iter(oneDegree * phiE_3 {;}% WE NEED TO HIDE THIS ; FROM deffloatvar <<< TO BE IMPROVED IN XINT
% initial value (in RADIANS) is E
% conversion of phi_3
subs((abs(D) < 1e-6)?
% this test of absolute error is not very good in floating point context
{break(@-D)}% stop now
{@ - D} % @ is previous value, replace it with @ - D
, D = f(@, a_e1, e_e1)/fD(@, a_e1, e_e1)
)% end of substitution of D value
, i = 1++)% iterate as many times as needed
* oneRadian;% end of definition of phiE (final value already converted to DEGREES)

% Now get the calculated angle phi_4 back from phiE
% We have defined E_to_phi_Pol as using floating point, hence the \xintfloatexpr wrapper
\xintdefvar phi_4 := \xintfloatexpr E_to_phi_Pol(phiE)\relax;

```

Faire le dessin :

```

\begin{pspicture}[showgrid=false,saveNodeCoors,NodeCoorPrefix=n](0,-1)(14,5)
\psset{dotscale=0.8}
\psset{PointSymbol=*,plotpoints=600}
\footnotesize
% Set nodes to center and foci
\pstGeonode[PosAngle=135,PointName={0},PointSymbol=+,dotscale=2,dotangle=45](\XcalcR{0x},\XcalcR{0y}){0}
\pnodes(\XcalcR{F_1x},\XcalcR{F_1y}){F_1}(\XcalcR{F_2x},\XcalcR{F_2y}){F_2}
\pstEllipseFocusNode[PosAngle=-90](0)(\XcalcR{a_e1},\XcalcR{b_e1}){F_1}{F_2}
% Set dots on the ellipse at the points with the polar angles phi_1, phi_2, ...
\xintforpair #1#2 in { (P_1,phi_1),(P_2,phi_2),(P_3,phi_3),(P_4,phi_4)}\do
{ \pnode(\XcalcR{[Ellipse(#2)][0]},\XcalcR{[Ellipse(#2)][1]}){#1}%
\psdot[dotstyle=*,dotscale=1]{#1}%
\uput[\XcalcR{trans_phi_M(#2)}]{#1}{#1$}
}
% Fill area between phi_1 and phi_2
\pscustom[fillstyle=solid,fillcolor=cyan,opacity=0.25,linestyle=none,polarplot,algebraic]{%
\translate(\XcalcR{F_2x},\XcalcR{F_2y})
\psplot{\XcalcR{phi_1*Pi/180}}{\XcalcR{phi_2*Pi/180}}{\XcalcR{b_e1^2/a_e1}/(1+\XcalcR{sqrt(a_e1^2-b_e1^2)/a_e1}*cos(x))}
\lineto(F_2)
\closepath
}
% Set the angle label
\pstMarkAngle[linecolor=cyan,MarkAngleRadius=0.8,LabelSep=0.5]{P_1}{F_2}{P_2}{\Delta\varphi_1$}
\psline[linecolor=cyan](F_2)(P_1)
\psline[linecolor=cyan](F_2)(P_2)
% Fill area between phi_1 and phi_2
\pscustom[fillstyle=solid,fillcolor=cyan,opacity=0.25,linestyle=none,polarplot,algebraic]{%
\translate(\XcalcR{F_2x},\XcalcR{F_2y})
\psplot{\XcalcR{phi_3*Pi/180}}{\XcalcR{phi_4*Pi/180}}{\XcalcR{b_e1^2/a_e1}/(1+\XcalcR{sqrt(a_e1^2-b_e1^2)/a_e1}*cos(x))}
\lineto(F_2)
\closepath
}
% Set the angle label
\pstMarkAngle[linecolor=cyan,MarkAngleRadius=1.8,LabelSep=1.4]{P_3}{F_2}{P_4}{\Delta\varphi_2$}
\psline[linecolor=cyan](F_2)(P_3)
\psline[linecolor=cyan](F_2)(P_4)
% orange dots at the foci
\psdots[dotstyle=o,dotscale=1,fillcolor=orange](F_1)(F_2)
% Draw ellipse and color parts of the ellipse
\pstEllipse[linecolor=black,linewidth=1pt](0)(\XcalcR{a_e1},\XcalcR{b_e1})
\rput(F_2){%
\psplot[polarplot,plotstyle=curve,linewidth=0.5pt,linecolor=Crimson]{%
{\XcalcR{phi_3}}{\XcalcR{phi_4}}{\XcalcR{b_e1^2/a_e1} 1 \XcalcR{sqrt(a_e1^2-b_e1^2)/a_e1} x cos mul add div}
}
}
\pstGeneralEllipse[linecolor=purple,linewidth=1pt](0)(\XcalcR{a_e1},\XcalcR{b_e1})%
[0][\XcalcR{phi_Pol_to_E(phi_1)}][\XcalcR{phi_Pol_to_E(phi_2)}]
\end{pspicture}

```